

Cosmology meets Quantum Gravity ?

Sergey Sibiryakov



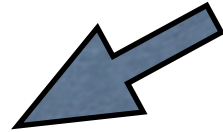
Benasque, 08/08/14

plethora of inflationary models
vs.
limited experimental information

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bottom-up



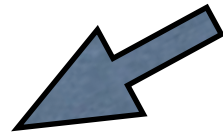
Effective theory of inflation

Cheung et al. (2008)

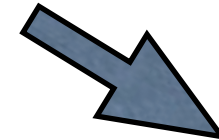
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bottom-up



top-down



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Postulate theoretical requirements to restrict the inflationary sector

An ideal theory of inflation:

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- As minimal as possible / falsifiable

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- Connected to particle physics
- Embedded into a UV complete theory including gravity

Examples of minimal models

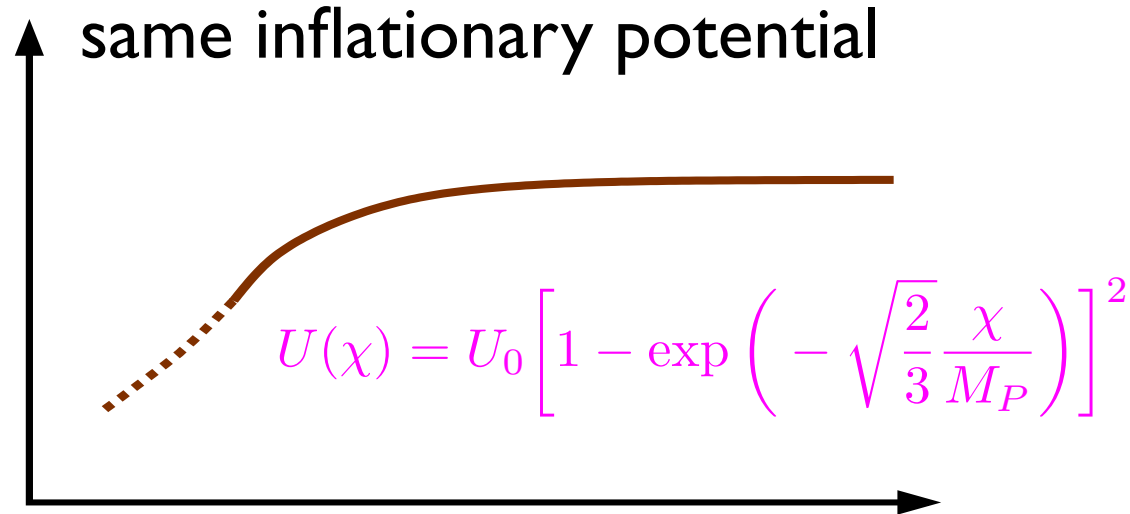
Starobinsky

$$-\frac{M_P^2 R}{2} + \frac{R^2}{6f_0^2} + \mathcal{L}_{SM}$$

Higgs

$$-\frac{M_P^2 R}{2} - \xi H^\dagger H R + \mathcal{L}_{SM}$$

same inflationary potential



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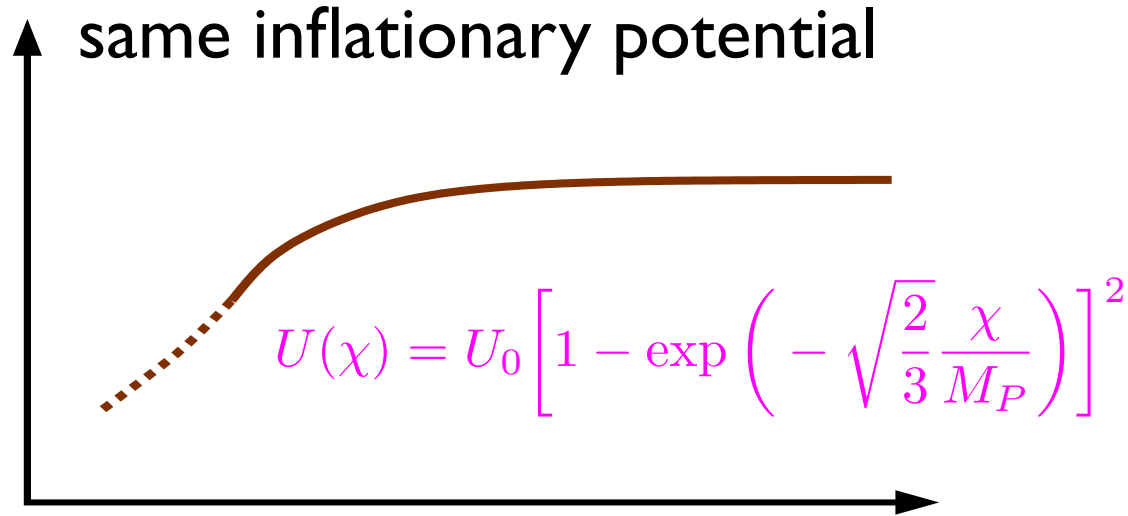
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but (slightly) different predictions because of different reheating

Bezrukov & Gorbunov (2011)

n_s	0.965	0.967
r	0.0036	0.0032

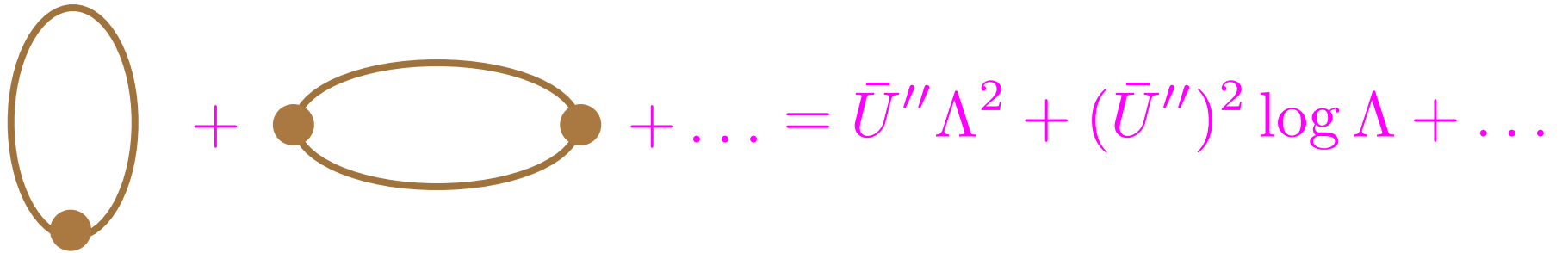
cf. CMBPol target $\delta n_s \approx 0.0016$, $\delta r \approx 0.5 \times 10^{-3}$

Beyond tree level ?

Bezrukov, Manin, Shaposhnikov, S.S. (2010)

George, Mooji, Postma (2014)

$$U(\bar{\chi} + \delta\chi) = U(\bar{\chi}) + \sum_{n=2}^{\infty} \frac{U^{(n)}(\bar{\chi})}{n!} (\delta\chi)^n$$



The diagrammatic expansion shows two Feynman diagrams. The first is a tadpole diagram consisting of a vertical loop with a single external vertex at the bottom. The second is a bubble diagram consisting of a horizontal loop with two external vertices on the left and right sides. These diagrams are followed by an ellipsis, then the mathematical expression $= \bar{U}'' \Lambda^2 + (\bar{U}'')^2 \log \Lambda + \dots$.

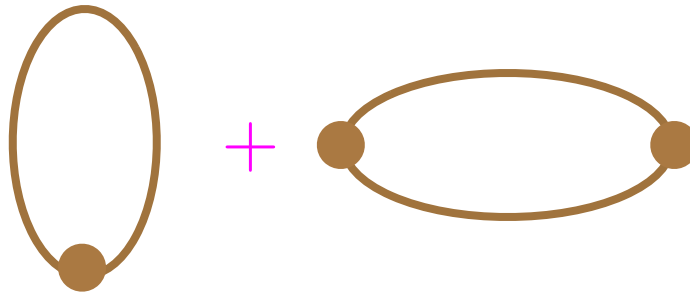
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The diagram shows two Feynman diagrams representing loop corrections. The first is a tadpole diagram consisting of a vertical loop with a single external vertex at the bottom. The second is a bubble diagram consisting of a horizontal loop with two external vertices at the left and right ends. Both diagrams are drawn in brown with solid circular vertices.

$$+ \dots = \bar{U}'' \Lambda^2 + (\bar{U}'')^2 \log \Lambda + \dots$$

At large χ the divergences are absorbed in an EFT-like expansion

$$U(\chi) = U_0 \left[1 + \sum_{n=1}^{\infty} u_n \exp \left(-n \sqrt{\frac{2}{3}} \frac{\chi}{M_P} \right) \right]$$

Physical principle: asymptotic shift symmetry at $\chi \rightarrow \infty$

Losses: no unique determination of slow-roll parameters

$$\eta \propto u_1, \quad \varepsilon \propto u_1^2$$

 **uncertainty of order 1** in $n_s - 1$

still robust prediction $\varepsilon = \frac{3}{4}\eta^2 + O(\eta^3)$

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e.g. the cutoff can be due to integration of a heavy field

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Seemingly, UV completion must involve gravity

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field theory in 4d
based on general
covariance

Quantum mechanics

unitary evolution in
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String theory

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- + encompasses particle physics and cosmology
able to produce concrete inflationary potentials

Silverstein & Westphal (2008)

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- very non-minimal
 - proliferation of extra fields
 - moduli stabilization
 - poorly understood in curved space-time
 - poorly understood with broken SUSY
 - no direct relation between inflation and low energies

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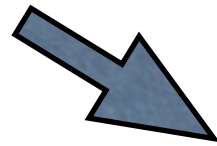
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Horava-Lifshitz gravity

Horava (2009)

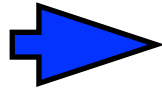
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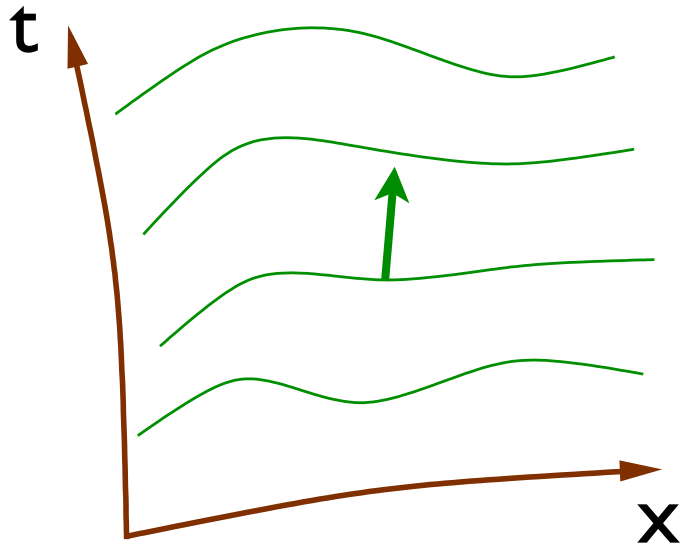
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$$\langle g_{\mu\nu} g_{\lambda\rho} \rangle \propto \frac{1}{\omega^2 - k^6}$$



loop integrals with gravitons are better behaved, the theory is “power counting renormalizable”



because of preferred foliation non-trivial effects persist at all energies

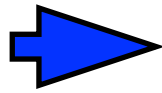
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e.g. $\alpha_1, \alpha_2 \propto (M_{QG}/M_P)^2$

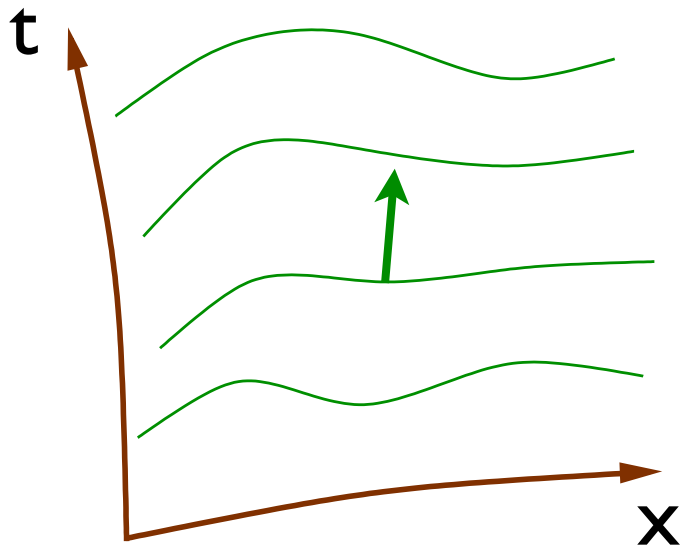
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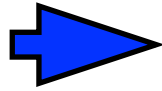
Constraints on QG from Solar System !

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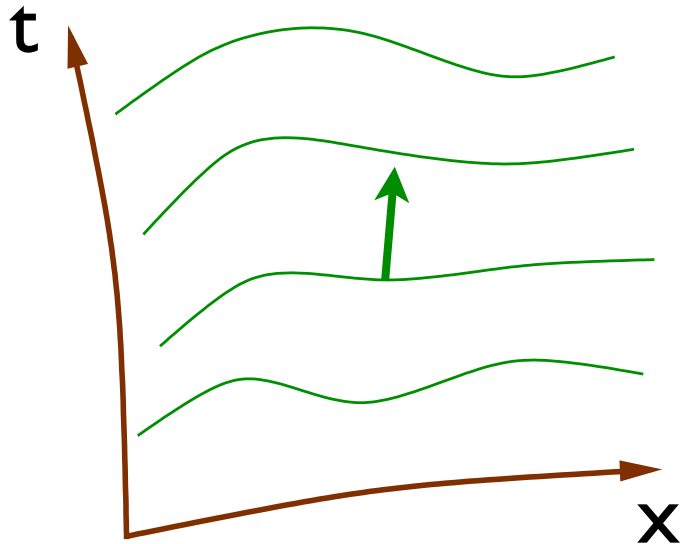
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More constraints from pulsars, CMB, LSS etc.

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Any $U(\chi)$ is renormalizable (divergences removed by normal ordering ?)

Naturally incorporates setups with asymptotic shift symmetry (the “Einstein-frame language”)

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Additional features: coupling to the preferred frame $m^2 u^\mu \partial_\mu \chi$

Donnelly, Jacobson (2010)

Solomon, Barrow (2014)

Ivanov, S.S. (2014)

during inflation kinetic energy of the inflaton can dominate over its potential energy --- “fast-roll”

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N.B. With $m \sim \text{eV}$ can serve as a model for DE

Blas, S.S. (2011)

Challenges

- Proof of renormalizability
gauge symmetry, instantaneous modes, technical complexity
- Emergence of Lorentz invariance in Standard Model

$$E^2 = m_i^2 + c_i^2 p^2 \quad |\delta c_p| \lesssim 10^{-22}$$

- Accidental symmetry due to 3d SUSY

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New particle physics nearby

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New particle physics nearby

Your theory is crazy, the question is whether it's crazy enough to be true.

Bohr to Pauli



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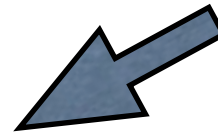
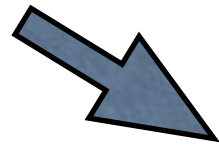
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Higher-derivative gravity

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Salvio, Strumia (2014)

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- Ghosts  loss of unitarity

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no unitary S-matrix at $E > M_P$

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Unification with the Standard Model and dark matter

Salvio, Strumia (2014)

Challenges

- Semiclassical interpretation

higher time-derivatives  spurious run-away solutions

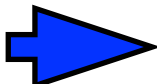
Proposal:

remove them by imposing boundary conditions in **future**

Maldacena (2011)

introduces a-causality; implications ?

- Good UV properties for $f_0^2 < 0$

$m_{\text{scalon}}^2 \propto f_0^2 M_P^2$  tachyonic instability

- stabilize by adding more fields ?
- accept run-away ?

Summary and Outlook



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 - to embed inflation in a broader context of particle physics
 - because observational precision approaches the level of quantum corrections
-  UV completion involves QG; some standard principles must be dropped
-  String theory seems too non-minimal to connect high and low-energy physics
-  Alternatives to ST drop either general covariance (and Lorentz invariance) or unitarity; many conceptual and technical challenges

potential gain: direct bridge between inflation and colliders