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# Gravity from Order and Number: Causal Sets

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June 2013

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## work with

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references

PI mini-course: http://pirsa.org/10100038 - 10100042

Joe Henson's "invitation to invitation": http://pirsa.org/10090092

Fay Dowker's public lecture to PI (available on You Tube and at PIRSA)

## I. What is QG? ("whole lot")

Among the various ideas put forward in the search for a theory of quantum gravity, the causal set hypothesis is distinguished by its logical simplicity and by the fact that it incorporates the assumption of an underlying spacetime discreteness organically and from the very beginning.

To put the causet programme in context, let me list some of the questions commonly asked in connection with quantum gravity.

#### What is the deep structure of spacetime? (kinematics)

continuous? discrete? nothing new? other?

(what kind of discreteness? crystalline, fluid)

how does this structure produce a continuum with metric? (cf. A puzzle)

why is there a notion of length at all?

(is the metric a field like any other?)

# What are the corresponding "deep laws of motion"?

in what sense need they be quantal?

what form should they take?

(Schrödinger eq? constraint equations? path integral form?)

(other form, eg 3rd order interference?)

causet answer: histories-based: soh of Schwinger-Kel'dysh type

These are dynamical questions. But will QG unify kinematics with dynamics?

Can we stick with the "Copenhagen interp" with its external observers, or do we need a better — more *intrinsic* — grasp of microscopic physics? Are Quantum Gravity and Quantum Foundations intertwined?

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# **Related theoretical questions**

How to describe non-gravitational matter (fermions, gauge-fields)? Can it arise spontaneously or must it be put in by hand? Is topology-change part of the theory? (eg to describe birth of cosmos, geon pairs) What becomes of unitarity and Relativistic Causality in QG? What are the "observables" of QG ("problem of time") Is the physical world "static" or "changing"?

#### Phenomenological questions

Why is  $\Lambda$  so small without being zero? (same as continuum problem!) Role of nonlocality? (radical nonlocality) Why d = 4? Why isn't spacetime full of holes? What is BH entropy counting? (horizon molecules?) And why the generalized second law? Does discreteness affect Hawking radiation? Are there observable signatures of discreteness?

Lorentz breaking?

New effects like swerves?

Initial conditions for cosmos?

why cosmos would have been so big and homogeneous at Planck time?

# II. Status of the causet programme

#### kinematics: good understanding of

dimension

length

homology

horizon area/entropy (both equilibrium and dynamical)

"problem of time" is solved

# dynamics in a background causet

classical (retarded) wave equation for scalar field QFT for scalar field

# dynamics of the causet

a classical dynamics of "sequential growth"

a "bilocal" action-functional S

in 2D we recover flat spacetime purely entropically

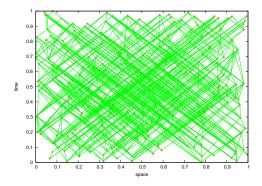
exploratory Monte Carlo simulations are being done for 2D, 4D

#### phenomenology

predicted  $\Lambda \sim 10^{-120}$  (batting 1000 (500?), 4 neutrinos??) toy model of Tolman-Boltzmann "cyclic" cosmology in CSG model of swerves: effect of discreteness on particle propagation (m = 0, m > 0) (bounds on phenomenological parameters) model of propagation of "scalar light" along causal links two proposals for black hole entropy, one based on the new action-functional *S* high frequency transparency (with caveat)

The recent progress on dynamics (of the causet and in a causet) stems from thinking about a scalar field in a causal set

# III. What is a causet?



define link

How does spacetime emerge?

geometry = order + number (N = V)

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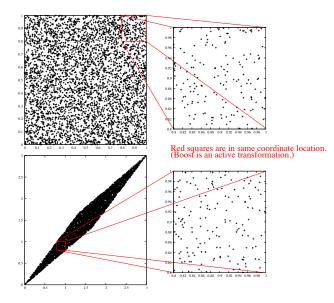
Meaning of  $M \approx C$  is not as obvious as for a fluid (-+++)

Role of sprinkling to obtain N = V (Poisson)

Sprinklings respect Lorentz symmetry (strictly)

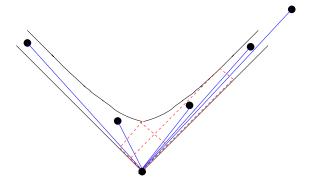
Theorem If *f* is a measurable equivariant partially defined mapping from Poisson sprinklings of  $\mathbb{M}^d$  to unit timelike vectors, then  $\operatorname{dom}(f)$  has measure zero.

In other words the probability that a sprinkling will break Lorentz symmetry is zero



Observe the radical nonlocality: numerous "nearest neighbors" (links)

This is why it's so hard to devise something like a local wave equation.



#### IV. The wave-equation and the causet action

Two roads to causet dynamics: principled and opportunistic The first has given us CSG but so far not QSG. So try the second.

Fix a causet  $C = \{e_1, e_2, e_3, ...\}$  that embeds faithfully in some region of  $\mathbb{M}^d$  via  $e_j \to x_j$ , and let  $\phi : \mathbb{M}^d \to \mathbb{R}$ . We seek a matrix  $\Box_{ik}$  such that  $\Box_{ik}\phi_k \leftrightarrow (\Box\phi)(x_i)$ .

The following works when d = 2 (overall scale-factor omitted):

 $\Box_{jj}=-1/2 \; ,$ 

 $\Box_{jk} = f(n_{kj}) \text{ when } e_j \succ e_k, \text{ where}$  $n_{kj} = \operatorname{card} \{ e_\ell | e_k \prec e_\ell \prec e_j \} \text{ and}$  $f(n) = 1, -2, 1, 0, 0, 0 \cdots \text{ [proof complete]}$ 

The same form works when d = 4 but with different coefficients:

 $\Box_{jj} = -1 \; ,$ 

 $f(n) = 1, -9, 16, -8, 0, 0 \cdots$  [proof almost complete]

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A 1-parameter generalization is also available ( $0 < \varepsilon \le 1$ ):

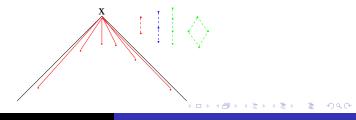
$$f(n) = \varepsilon (1-\varepsilon)^n \left[ 1 - 9 \frac{n\varepsilon}{(1-\varepsilon)} + 8 \frac{n(n-1)\varepsilon^2}{(1-\varepsilon)^2} - \frac{4n(n-1)(n-2)\varepsilon^3}{3(1-\varepsilon)^3} \right]$$

In curved spacetime our "D'Alembertian" becomes  $\Box - \frac{1}{2}R$  (Benincasa) Applying it to  $\phi \equiv -1$  we obtain

$$rac{1}{2}\int RdV \leftrightarrow -\sum_{jk}f(n_{jk})$$

For  $\varepsilon = 1$  this is just a weighted sum over a finite number of "building blocks", eg the diamond order-interval has weight f(2) = 16 for d = 4.

(There even seems to be a kind of Gauss-Bonnet result for d = 2.)



# What is the physical significance of Scauset?

Even in continuum no one really knows how  $S_{grav}$  is to be used — not even formally!

Nevertheless, we can attempt Monte Carlo

(Many questions arise: What to "measure"? To Wick rotate the causet itself makes no sense. We can nevertheless rotate S (each term has a definite sign). Might **quantum** montecarlo handle complex amplitudes better?)

Does S also signify an entropy somehow: Fay Dowker's "mutual information"!

## V. Quantum field theory à la causet

Recall that the continuum theory uses

both  $\Box \phi = 0$  and  $[\hat{\phi}(x), \hat{\phi}(y)] = i\Delta(x, y)$ 

where  $\Delta(x, y) = G(x, y) - G(y, x)$  and  $G = G^{retarded}$ .

These are compatible since  $\Box \Delta = 0$ 

On the causet they are not compatible, so use the commutator only (Noldus)

Can the commutator alone yield the full algebra of operators (and the "vacuum state")?

For a Gaussian (Wickian) field, everything follows from  $W(x, y) = \langle \hat{\phi}(x) \hat{\phi}(y) \rangle$ (we assume also  $\langle \hat{\phi} \rangle = 0$ )

The antisymmetric part of W is just  $i\Delta/2$ 

Choose the symmetric part to make  $W \ge 0$  (Johnston)

$$2W = i\Delta + \sqrt{-\Delta^2}$$

We did this without ever being asked to define positive frequency! Byproduct is a unique "ground state" for any compact region of spacetime

All these steps can be repeated in the causet once we have  $G_{ik}^{retarded}$ .

In d = 2 we know G in full generality In flat d = 4 we also know G for all  $m^2$ In general we might try inverting the matrix  $\Box_{ik}$ 

Lots more to say: decoherence functional, interactions, classical limit, "causality" .....

Byproduct: new formula for entropy direct from  $W(x, y) = \langle \phi(x)\phi(y) \rangle$ 

With it we can define entanglement entropy of black hole in a causet ....