Cosmic Strings Stabilized

by Fermion Fluctuations

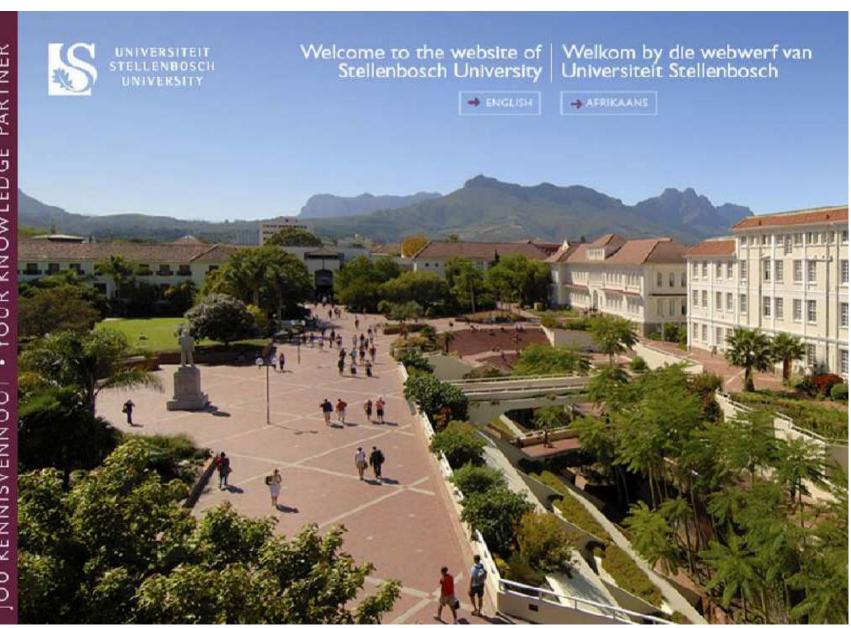
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Presentation is based on:

HW, M. Quandt, N. Graham, and O. Schröder, Nucl. Phys. B831 (2010) 306 HW and M. Quandt, Phys. Lett. B690 (2010) 514
HW, N. Graham, and M. Quandt, Phys. Rev. Lett. 106 (2011) 101601
N. Graham, M. Quandt, and HW, Phys. Rev. D84 (2011) 025017





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can be perfectly studied with spectral methods

- \star fermion bound state and vacuum energies are same $\mathcal{O}(\hbar)$
- \star non–trivial structure at spatial infinity
- \star gauge invariant field combinations are smooth at spatial infinity
- \star individual Feynman diagrams are not gauge invariant

* static string profiles in a $SU_L(2)$ gauge theory

• vector meson (temporal gauge):

$$\vec{W} = n\sin(\xi_1)\frac{f_G(\rho)}{g\rho}\hat{\varphi}\begin{pmatrix}\sin(\xi_1) & i\cos(\xi_1)\,\mathrm{e}^{-in\varphi}\\ -i\cos(\xi_1)\,\mathrm{e}^{in\varphi} & -\sin(\xi_1)\end{pmatrix}$$

• Higgs meson:

$$\Phi = v f_H(\rho) \begin{pmatrix} \sin(\xi_1) e^{-in\varphi} & -i\cos(\xi_1) \\ -i\cos(\xi_1) & \sin(\xi_1) e^{in\varphi} \end{pmatrix}$$

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$$\begin{array}{ll} \rho \longrightarrow 0: & f_G, f_H \longrightarrow 0\\ \rho \longrightarrow \infty: & f_G, f_H \longrightarrow 1 \end{array}$$

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- *massless* particles at the origin, gain of energy?
- non-trivial structure at spatial infinity: neither Born nor Feynman series well defined

 \star phase shift approach: summing the changes of zero–point energies

$$E_{\rm vac} = \frac{\hbar}{2} \sum_{n} \left(\omega_n - \omega_n^{(0)} \right) \bigg|_{\rm ren.} = \frac{\hbar}{2} \sum_{j} \epsilon_j + \hbar \int dk \, \omega_k \, \Delta \, \rho_{\rm ren.}(k)$$

 ϵ_j : true bound state energies

 $\omega_k = \sqrt{k^2 + m^2}$ energy of continuum states

- $\Delta
 ho(k)$: change in density of continuum states (in momentum space) = $\frac{1}{\pi} \frac{d}{dk} \delta(k)$
 - $\delta(k)$: phase shift

II) Spectral (Phase Shift) Method

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 \star requires angular momentum decomposition

\star interface formalism (energy per unit length)

- start from energy momentum tensor in field theory
- integrate over momentum of translationally invariant subspace
- dimensional regularization for associated UV divergence (*)
- scattering problem in orthogonal two-dim. subspace
- cancel divergence (*) by sum rules for scattering data generalization to Levison's theorem

$$E_{\delta}^{(N)} = \frac{1}{4\pi} \sum_{\ell} \left\{ D_{\ell} \int_{0}^{\infty} \frac{dk}{\pi} \left[\omega_{k}^{2} \ln \left(\frac{\omega_{k}^{2}}{\mu^{2}} \right) - k^{2} \right] \frac{d}{dk} [\delta_{\ell}(k)]_{N} + \sum_{j} \left[(\epsilon_{j,\ell})^{2} \ln \frac{(\epsilon_{j,\ell})^{2}}{\mu^{2}} - (\epsilon_{j,\ell})^{2} + m^{2} \right] \right\}$$

integration before summation!

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• exchange of sum and integral only possible after rotating to imaginary momentum axis, t = ik.

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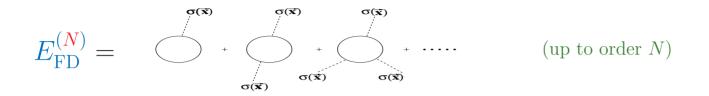
• δ_{ℓ} : total phase shift in (angular momentum) channel ℓ

(phase of detS)

- $\omega_k = \sqrt{k^2 + m^2}$: energy of continuum mode
- D_{ℓ} : degeneracy in (angular momentum) channel ℓ
- N: number of Born subtractions to render integral finite
- $\epsilon_{j,\ell}$: bound state energies
- μ : redundant renormalization scale

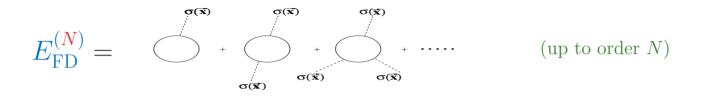
(sum rules for scattering data)

 \star Feynman diagrams (add back in subtractions)



• σ : background potential induced by string

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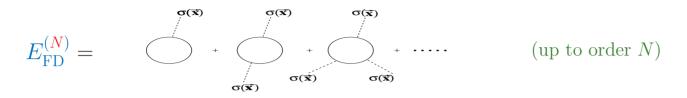
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 \star counterterms

$$E_{\rm CT} = \sum_{i} c_i E^{(i)}[\sigma]$$

- c_i : counterterm coefficients computed from renormalization condition in the perturbative sector
- $E^{(i)}$: local (gauge invariant) energy functionals from the classical theory

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 \star total vacuum energy

$$E_{\rm vac} = E_{\delta}^{(N)} + E_{\rm FD}^{(N)} + E_{\rm CT}$$

III) Fermions in String Background

 \star fermions dominate vacuum energy at large N_c \star pure gauge at spatial infinity

 $U = P_L \exp\left(i\hat{n} \cdot \vec{\tau}\xi_1\right) + P_R \quad \text{with} \quad \hat{n} = \begin{pmatrix} \cos(n\varphi) \\ -\sin(n\varphi) \\ 0 \end{pmatrix}$ so that $g\vec{W} \sim U^{\dagger}\nabla U \quad \text{and} \quad \Phi \sim U\Phi_0$ \star pure gauge at spatial infinity

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 \star return string at spatial infinity (as for QED flux tubes)

- numerically very costly
- proof–of–principle calculation:

vacuum polarization energy small in $\overline{\text{MS}}$.

 \star major differences to QED flux tubes:

• gauge transformation that unwinds the string is unique:

 $U(\varphi) = U(\varphi + 2\pi)$

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 \star radially dependent gauge transformation

 $H \longrightarrow U^{\dagger}(\rho, \varphi) H U(\rho, \varphi)$

with

$$U(\rho,\varphi) = P_L \exp\left(i\hat{n} \cdot \vec{\tau}\,\xi(\rho)\right) + P_R$$

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 \star local gauge invariance: specific shape of ξ irrelevant

 \longrightarrow further support for phase shift method

 \star field parameterization

$$f_H(\rho) = 1 - e^{-\frac{\rho}{w_H}}$$
$$f_G(\rho) = 1 - e^{-\left(\frac{\rho}{w_G}\right)^2}$$
$$\xi(\rho) = \xi_1 \left[1 - e^{-\left(\frac{\rho}{w_\xi}\right)^2}\right]$$

- w_H , w_G , ξ_1 : variational parameters for physical string
- w_{ξ} : gauge parameter, not measurable
- similarly, the scale parameter for the fake boson field is not measurable (technicality to simplify 3rd and 4th order FDs)

 \star data in $\overline{\text{MS}}$ scheme (scale set by fermion mass)

w_{ξ}	$E_{\rm FD}$	E_{δ}	$E_{ m B}$	$E_{ m vac}$	
0.5	-0.2515	0.3489	0.0046	0.1020	
1.0	-0.0655	0.1606	0.0032	0.0983	$w_H = 2.0$
2.0	-0.0358	0.1294	0.0038	0.0974	$w_G = 2.0$
3.0	-0.0320	0.1235	0.0056	0.0971	$\xi_1 = 0.4\pi$
4.0	-0.0302	0.1193	0.0080	0.0971	

- $\bullet~E_{\rm FD}$ renormalized first and second order Feynman diagram
- E_{δ} phase shift contribution, first and second Born order removed
 - \ast computed by analytic continuation
 - \ast log. div. of 3rd and 4th order by a fake field
 - \ast numerically very costly
 - * about $1 \dots 2\%$ numerical error
- $E_{\rm B}$ remnant of fake field

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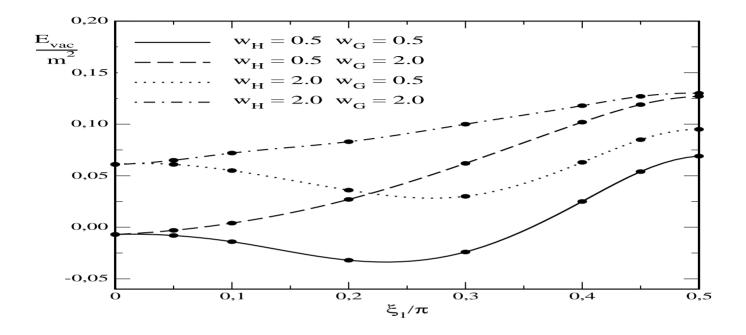
spectral methods to compute vacuum energies verified

 \star on–shell scheme for $E_{\rm CT}$

add (maximally four) finite local and gauge-invariant counterterms:

- no-tadpole interaction
- Higgs mass unchanged
- normalization of Higgs particle unchanged
- normalization of vector meson unchanged
- vector meson mass will be a prediction,
 - \longrightarrow tune gauge coupling to reproduce M_W

\star results in on–shell scheme



- finite counterterm contribution adds positively
- quantum effects provide some binding for thin strings (Landau ghost ?)
- competitive with classical mass only for
 - (i) fermion masses about 1TeV
 - (ii) many internal degrees of freedom (color)

- \star total energy: $E = E_{\rm cl} + 3E_{\rm vac}$ $(N_C = 3)$
- \star classical contribution:

$$\frac{E_{\rm cl}}{m^2} = 2\pi \int_0^\infty \rho d\rho \left\{ n^2 \sin^2 \xi_1 \left[\frac{2}{g^2} \left(\frac{f'_G}{\rho} \right)^2 + \frac{f_H^2}{f^2 \rho^2} \left(1 - f_G \right)^2 \right] + \frac{f'_H^2}{f^2} + \frac{\mu_h^2}{4f^2} \left(1 - f_H^2 \right)^2 \right\}$$

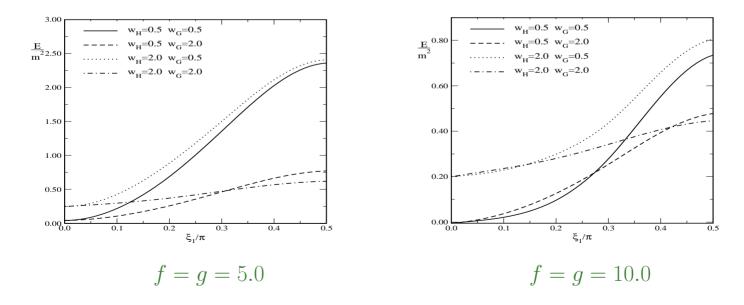
(quantities under the integral are scaled to be dimensionless, $\mu_h = \frac{m_H}{vf}$)

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\star numerical results:



IV) Charged Cosmic Strings

 \star Cosmic strings induce fermionic bound state levels: $\epsilon_i < m$

- ϵ_i eigenvalues of two–dimensional Hamiltonian
- for $\xi_1 = \pi/2$ an exact zero mode exists
- for wide strings, $w_{\rm H} \gtrsim 4$ many levels emerge (100)

 \star bound states carry longitudinal momenta:

$$E_i(p_n) = \sqrt{\epsilon_i^2 + p_n^2} \qquad p_n = \frac{n\pi}{L}$$
 (length of the string)

 \star populating these levels may form a charged string with energy less than equally many free fermions

* binding energy is of same order as E_{vac} , both in N_{C} and \hbar . * require leading contribution to the total energy as $L \to \infty$

$$\sum_{n} \longrightarrow \frac{L}{\pi} \int dp$$

 \star search for the minimum of bound state contribution

- introduce chemical potential $\mu \leq m$
- populate all levels with $E_i(p) \le \mu$
- Fermi momentum for each populated level $p_i^F(\mu) = \sqrt{\mu^2 \epsilon_i^2}$ (states with $\epsilon_i > \mu$ are not populated)

 \star charge density (per unit length)

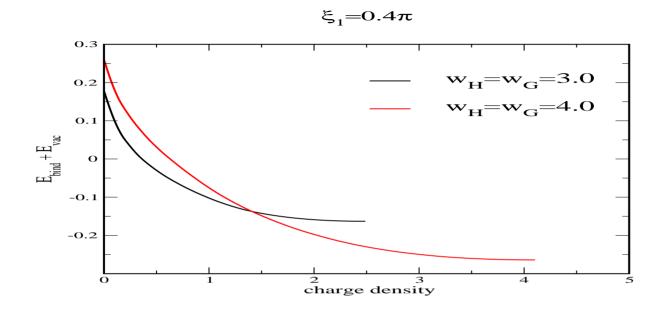
$$Q(\mu) = \sum_{i} \frac{p_i^F(\mu)}{\pi}$$

monotonously rising, can be inverted

$$\mu = \mu(Q) \qquad \text{and} \qquad p_i^F = p_i^F(Q)$$

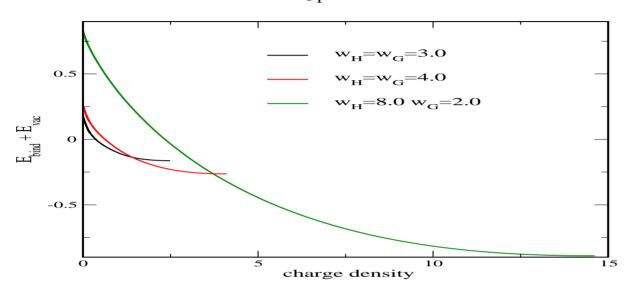
 \star binding energy (density) at prescribed charge (density)

$$E_{\text{bind}}(Q) = \frac{1}{\pi} \sum_{i} \int_{0}^{p_{i}^{F}(Q)} dp \left[\sqrt{\epsilon_{i}^{2} + p^{2}} - m \right] \qquad \left(\begin{array}{c} \text{relative to} \\ \text{free fermions} \end{array} \right)$$

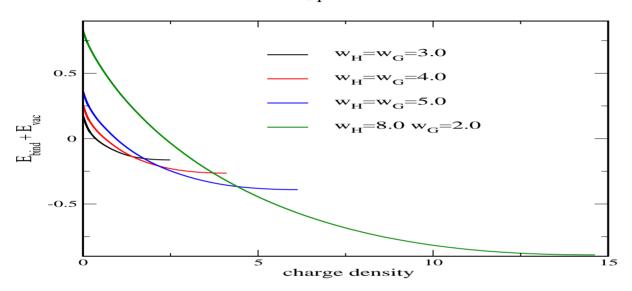


- endpoint: all available levels populated
- small charge: narrow string preferred
- large charge: wide string preferred

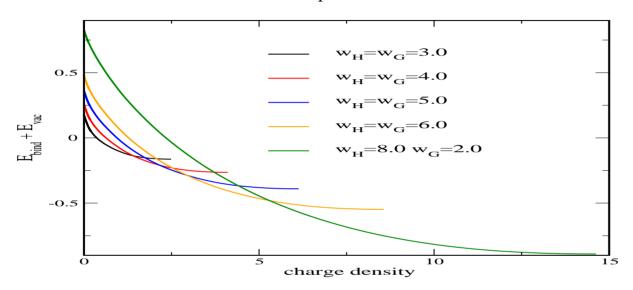
\star adding more configurations



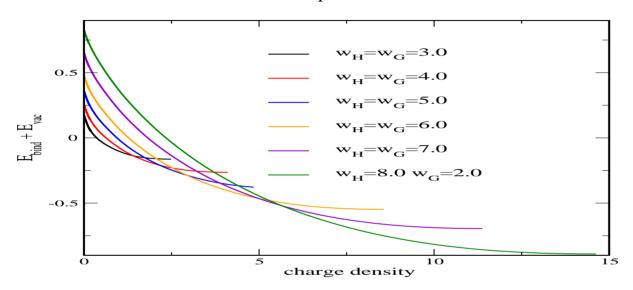
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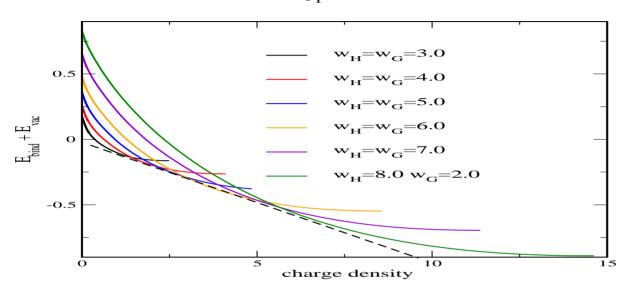
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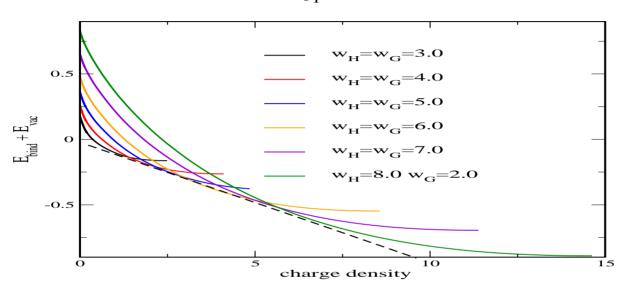


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 $\min\left(E_{\rm bind} + E_{\rm vac}\right) \propto Q$

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- essentially independent of w_G and ξ_1



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 \star identify configurations along the envelope to compute associated $E_{\rm cl}$

- \star about $m = 1, \ldots, 500$ configurations for which E_{vac} and ϵ_i have been calculated
- \star prescribe Q and compute

$$E_{\rm tot}^{(m)}(Q) = E_{\rm cl}^{(m)} + N_C \left[E_{\rm vac}^{(m)} + E_{\rm bind}^{(m)}(Q) \right]$$

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 \star negative search for physically motivated parameters (top–quark)

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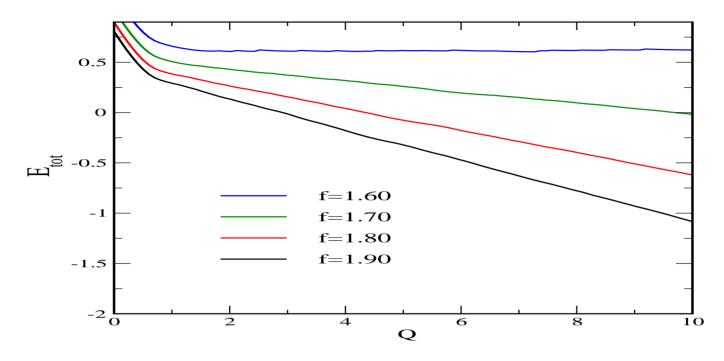
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 \star but: $E_{\rm cl}^{(m)}$ decreases rapidly with increasing f



 \star stable configurations for $f \gtrsim 1.7$

- \star safe against Landau ghost problems (widths parameters large)
- \star stable strings should emerge if a strongly interacting fermion doublet with twice the top quark mass existed
- * stable configurations have $\xi_1 \approx 0$ (since we keep g small)

\star total mass

- \bullet small to moderate binding: $M\approx LQm$
- typical charge (density) for binding: $Q\approx 5m$
- typical fermion mass at binding: $m \approx 300 \text{GeV}$
- cosmological scale: $L \approx R_{\odot}$

 $M \approx 10^{-20} M_{\odot}$

• too light to have cosmological impact

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- \star future work: boson loops
 - currents along the string
 - closed strings